

# Phase stability and thermal equations of state of Fe<sub>3</sub>S and Fe<sub>2</sub>S polymorphs to Earth's core pressures and high temperatures

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## Abstract

The seismologically-determined density profile and dynamics of Earth's iron-rich core can only be explained by incorporating a small but significant light element component such as sulfur. Understanding the phase stability and pressure-temperature-volume relationships of iron-rich sulfides to core conditions is critical for assessing the core's composition. In this *in-situ* X-ray diffraction study, we determined the high *P-T* stability fields of Fe<sub>3</sub>S and Fe<sub>2</sub>S polymorphs to outer core pressures and high temperatures and fit their thermal equations of state. Between 26 and 140 GPa and at moderate temperatures, a *Pnma* Fe<sub>2</sub>S structure coexists with Fe and FeO. Upon heating, below 85 GPa, the *Pnma* Fe<sub>2</sub>S phase transitions to  $\bar{I}4$  Fe<sub>3</sub>S. Above 85 GPa, the *Pnma* Fe<sub>2</sub>S phase transitions to a  $P\bar{6}2m$  Fe<sub>2</sub>S structure at high temperatures. This phase is stable on the liquidus to 142 GPa. We fit thermal equations of state for  $\bar{I}4$  Fe<sub>3</sub>S and *Pnma* Fe<sub>2</sub>S to 75 GPa and 137 GPa respectively and determined the following isothermal bulk moduli ( $K_0$ ) and pressure derivatives ( $K_0'$ ):  $K_0 = 124 \pm 2$  GPa and  $K_0' = 5.2 \pm 0.1$  for Fe<sub>3</sub>S and  $K_0 = 149 \pm 11$  GPa and  $K_0' = 5.1 \pm 0.3$  for Fe<sub>2</sub>S. Fe<sub>3</sub>S was found to compress isotropically while the *Pnma* Fe<sub>2</sub>S phase displayed anisotropic compressibility. Extrapolating the adiabatic density curves of Fe<sub>3</sub>S and *Pnma* Fe<sub>2</sub>S to outer core conditions, we concluded that 15-18 wt% S is required to account for the density deficit at the core-mantle boundary. Based on the phase relations determined in this study, Fe<sub>2</sub>S, not Fe<sub>3</sub>S is stable at outer core pressures and high temperatures and could play a critical role in the thermodynamics of Earth's core.